

## 2017 Midterm Review

### Galaxy classification – BM Ch 4.1

Hubble and other classification systems

Trends of physical (gas content, star formation) and morphological galaxy properties along Hubble sequence

### Photometric Properties of Ellipticals – BM 4.2-4.3, SG 6.1

Surface Brightness profile for E's – DeVaucouleurs ( $R^{1/4}$ ), Sersic

Centers of Ellipticals (core/coreless)

Triaxiality – Disky/boxy isophotes and relation to other physical properties

### Photometric Properties of Spirals – BM 4.4, SG 5.1

Surface Brightness profile for disks - exponential

vertical distribution – scale height

### Ellipticals – stellar pops and ISM - BM 4.3, 8.3, 11.1, 11.2, SG 6.3

Trends with luminosity, color, metallicity, gas content, X-ray/radio emission

Kinematics – how do we measure velocity dispersion and rotational velocity

Typical values for  $V_{\text{rot}}$  and sigma

Faber-Jackson ( $L \sim \text{sigma}^4$ )

Fundamental plane and how it relates to VT

### Spirals – ISM, structure and Kinematics - BM 8.2, 11.3, SG 5.2-5.3

Gas content/distribution, velocity maps, rotation curves

MW and the Sun's LSR (know  $V_o$  and  $R_o$ )

Typical values for  $V_{\text{rot}}$  and sigma

Tully-Fisher ( $L \sim V_{\text{max}}^4$ )

$$M(R) = \frac{V^2(R)R}{G} \rightarrow \rho(R) = \frac{V^2}{4\pi GR^2}$$

Metallicity in disks (compared to closed-box chemical evolution model)

Spiral structure - density waves, kinematic spiral

### Potential Theory and Orbits - BT 2.0-2.2, 3.0-3.4, SG 3

Gravitational potentials – Poisson's equation  $\nabla^2\Phi(R, z) = 4\pi G\rho(R, z) = \frac{\partial^2\Phi}{\partial z^2} + \frac{1}{R} \frac{\partial}{\partial R} \left[ R \frac{\partial\Phi}{\partial R} \right]$

Spherical potentials and orbits – rosettes

Integrals of motion – conservation of energy and momentum

$$\Phi = -\frac{GM}{R} \quad \Phi = \frac{1}{2}\Omega^2 R^2 + const$$

Two extreme potentials

Axisymmetric potentials and orbits  $\Phi_{eff} = \Phi(R, z) + \frac{L^2}{2R^2}$

Nearly Circular Orbits – epicycle approximation

Crossing time/Relaxation time  $t_r = \frac{V_v^2}{\Delta V^2} t_c \simeq \frac{N}{8 \ln N} t_c$

Collisionless Boltzman Equation

**Example question:**

What is the fundamental plane? How does it relate to the virial theorem and what does this relation tell us about the type of galaxies that lie in the fundamental plane?

Draw a typical rotation curve for a spiral galaxy. Derive the density distribution required to produce this rotation curve at large radii. How does this indicate the presence of dark matter?

Calculate the mass interior to the Sun's orbit using accepted values for  $V_0$  and  $R_0$ .